Fast stars challenge big bang origin for dwarf galaxies

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Our Galaxy is a member of a group of at least 30 galaxies called the Local Group, which extends more than three million light years across space. Over half are “small”, or dwarfs, less than 6,000 light years in diameter. These dwarf galaxies contain stars that are moving away from each other at high velocities, as shown by their measured radial velocities.

Stars have what is known as space velocity, or movement in space with respect to our sun. Space velocity has two components: (i) radial velocity (toward or away from us) and (ii) proper motion (at right angles to our line of sight).

Proper motion is measured by comparing the position of stars on two different photographic images of the same region of the sky taken at least a decade apart. However the Local Group dwarf galaxies are too far away for us to observe changes in proper motion, so we can measure only radial velocities.

The radial velocities for the dwarfs are obtained by analysing the spectra of various stars found in them. Astronomers assume that the Doppler shift applies and interpret blue shifts as indicating motion toward the earth and red shifts indicating motion away from the earth. If, as some have challenged (e.g. Halton Arp, Barry Setterfield), redshifts are not the result of motion away from Earth, or 3D space expanding as in the big bang cosmology, this could significantly affect the interpretation of radial velocity measurements commonly used today in astronomy. But the Doppler interpretation of spectral shifts seems to be valid for the stars in dwarf galaxies.

After astronomers have measured the radial velocity of a number of stars in a dwarf galaxy, they can work out how fast the stars are moving relative to each other. Surprisingly, many stars in the dwarf galaxies in the Local Group are moving away from each other at speeds close to 10–12 km/s.

Astronomers recognise that, at face value, these high velocity dispersion stars indicate that they have not been moving through the dwarfs for billions of years. These stars could pose a major problem to big bang cosmology and the Hubble age of the universe.

For astronomers who are committed to the big bang model and a universe billions of years old, these high velocity stars are a major problem. At these speeds, the stars should have dispersed in 100 Ma, which, compared with the supposed 15,000 Ma age of the universe, is a short time. How is it possible that the dwarf galaxies still contain such speedy stars? The answer, we are told, is ‘dark matter’.

The only way that the stars could still be held inside the dwarf galaxies and not have escaped billions of years ago is if there is something holding them together. The proposal is that the dwarf galaxy contains a huge amount of additional mass that, except for its gravitational force, is entirely invisible and undetectable. This simple approach is based upon Newton’s 1st Law of Motion. Using this law and a dispersion velocity of 12 km/s, the stars would travel about 4,000 light years from their homes in only 100 Ma. The dark matter needs to surround and permeate the dwarfs to retain the speedy stars for billions of years.

Astronomers researching the Local Group dwarfs use the M/L (mass-to-light) ratio, where both parameters are quoted in solar units (for the Sun, M/L = 1.0). The M/L ratio is an interpretation based upon stellar velocities measured in the galaxies, assuming that the galaxies still exist after billions of years because they are gravitationally bound. The higher the mass-to-light-ratio, the greater the amount of dark matter needed to keep the galaxy together. However, the high M/L ratios could be used to support a very young dynamical age for the Local Group dwarfs instead. Perhaps dwarf galaxies that exhibit large M/L ratios with
speedy stars may simply be evidence for a young age and not objects that are billions of years old. There may be no need for a large quantity of dark matter, since a young universe does not necessarily require galaxies (or clusters of galaxies) to be gravitationally bound.

Astronomer Ken Croswell (Ph.D. in astronomy from Harvard University, who contributes articles from time to time in New Scientist, Astronomy, and Sky & Telescope) clearly recognises the problem that high velocity dispersion stars present for our neighbouring dwarf galaxies like Draco and Ursa Minor. Keep in mind that Draco is only about 4,500 light-years in diameter while Ursa Minor is even smaller at 3,000 light-years. 'But for a tiny galaxy like Draco, it [the speedy star problem] is enormous. To hold on to such speedy stars, dim Draco must have a large mass and an enormous quantity of dark matter.'

'The stars of Ursa Minor and Draco have velocity dispersions of around 12 kilometres per second. If these stars are not gravitationally bound to their parent galaxies, a velocity of 12 kilometres per second will carry the stars 4000 light-years from their homes in only 100 million years, meaning that a dwarf would vanish soon after it fell victim to the Galactic tide.'

The Galactic tide is a differential gravitational force exerted by the Milky Way on the Local Group dwarf galaxies. The edge of a dwarf nearest the centre of the Milky Way feels a stronger pull than the edge farther away. In theory, dwarf galaxies could be torn apart and lose stars, especially the high velocity dispersion stars.

It is clear that high velocity dispersion stars should have escaped from the dwarfs long ago according to big bang cosmology and Hubble time. So, to save their theory, big bang astronomers invoke a number of difficult-to-test and perhaps incapable-of-proof hypotheses, like dark matter.

Just what is dark matter? In the evolution model, dark matter has been thrown around with several definitions. Some believe it is not normal matter (composed of protons and neutrons), but some kind of hypothetical subatomic particles. These fall into two groups, hot dark matter for the hypothetical subatomic particles that zip around near light speed (best candidate is neutrinos if they have mass) and cold dark matter that travels at much slower speeds. Others believe dark matter is not exotic but perhaps brown dwarfs, balls of gas much larger than Jupiter but not large enough to ignite and sustain hydrogen fusion. Still more recent reports claim that dark matter may be nothing other than ordinary molecular hydrogen (H₂).

Is there any dark matter in the solar system? No. There was a report on this topic in the August 1, 1995 Astrophysical Journal submitted by a team from JPL, the University of Texas at Austin, and Southern Methodist University in Dallas. The report ends with the conclusions of Duane Dicus from the University of Texas:

'We have no evidence for dark matter that is trapped around the Sun.'

To be fair, evolutionists would not expect to find a significant amount of dark matter in the solar system or the local neighbourhood from the galactic rotation profile. Orbital velocities of stars in the inner portions of spiral galaxies (including the Milky Way) can be accounted for by the mass of the observed stars (and gas and dust) alone. But the outer portions of spiral galaxies require a substantial amount of dark matter to ensure they are gravitationally bound.

Another recent effort to detect dark matter in the solar vicinity derives from the Hipparcos data. Motions of spectral class-A stars were examined out to a distance of about 125 parsecs (410 light-years) from Earth.

'This sample provides the first reliable, homogenous tracer of mass and stellar motions in our part of the galaxy. It indicates that the total mass in the solar neighborhood amounts to 0.076 ± 0.015 solar mass per cubic parsec, roughly twice the amount from stars alone and well below all previous determinations — though compatible with all observations of interstellar gas and dust. This result leaves no room for significant dark matter in the Milky Way’s disk. A nearly spherical halo of dark matter extending to the galaxy’s outer reaches is still needed, however, to explain the high speeds at which the Milky Way’s outlying
Again we see the issue of a high M/L ratio when applied to the rate of revolution for stars, dust, and gas in the Milky Way’s outer edges. Whatever dark matter may be in the evolution model, it certainly is not measurable in our solar system or within the solar vicinity in interstellar space. I prefer astronomer Jeffrey Kuhn’s skeptical position on the dark matter issue as related by Ken Croswell:

“I have rather heretical views about the dark matter question. I keep thinking about the luminiferous ether from years past. Over a century ago, physicists convinced themselves that space was pervaded by an invisible substance called the ether, through which light waves propagated. Perhaps, said Kuhn, dark matter is today’s equivalent of the ether and does not exist.”

This suggestion, that dark matter may not be real, is worth pursuing. What we observe in our neighbouring galaxies, may point to a young universe or, at least, a recent origin for the Local Group. On-going research continues to show that the Local Group dwarfs need a significant amount of dark matter if they are to gravitationally hold the speedy stars for billions of years.

I consider that the high velocity dispersion stars and the interpretation of the high M/L ratios in dwarf galaxies might provide good evidence in the Local Group for an abrupt and recent origin of the universe or at least the Local Group. These fast stars challenge the big bang cosmology and the Hubble time for the age of the universe. It should be noted that:

a. The creation model does not need to invoke dark matter to explain cosmology or the origin of the universe. If dark matter were eventually confirmed and clearly defined as to what it may be, the creation model would neither gain nor lose. However, if future observations show that dark matter is insufficient to bind galaxies (or clusters thereof), this would be very strong evidence supporting a young universe.

b. The evolution model is very dependent upon the reality of dark matter to support a billion-year-old universe. Evolutionists invoke dark matter to explain how the first Population III stars formed, how galaxy clusters have been held together, and how speedy stars have remained inside the Local Group dwarf galaxies.

c. A young universe does not require the dust, gas, and stars in spiral galaxies to be gravitationally bound to complete multiple revolutions over billions of years. Perhaps the galaxies are slowly dispersing and will not complete multiple revolutions. In the same way, the speedy stars in Local Group dwarf galaxies may be slowly escaping.

The issue of dark matter and radial velocities of speedy stars in the Local Group dwarfs was reported in 1994 in popular science sources. At a conference at the Observatoire de Haute Provence in France, Mario Mateo presented an overview of the spectral data obtained on the radial velocities of the speedy stars in the dwarf galaxies. The consensus was that the dwarfs needed dark matter halos to hold the stars in the galaxy. Without halos of dark matter they would disperse into intergalactic space.

The 1994 reports described three observed features of the Local Group of dwarf galaxies that pose serious problems for the billions-of-years evolutionary scenario and that are under active investigation:

- Efforts to extrapolate orbital interactions between members of the Local Group over billions and billions of years.
- ‘Young’ Population I stars found along with ‘old’ Population II stars observed in the dwarf galaxies.
- The amount of dark matter contained in them to retain the speedy stars.

All these observations may support the concept of a recent creation featuring an abrupt origin of stellar bodies.

References


11. Hodge, P., Ref. 10, p. 29.